

Exam 3

Calculators are not allowed for this exam.

You will find that for many questions it is useful to know that $e^{-1} \approx 0.4$ and that other powers can be obtained using $e^{-n} = (e^{-1})^n \approx (0.4)^n$.

1. This exam contains 9 pages of questions and instructions, three pages of equations, a page of wavefunctions, a periodic table, and a page of constants and conversion factors.
2. Show your work and make your reasoning clear.
3. You have 2.0 hours to work on the exam.

1. _____/20

2. _____/25

3. _____/15

4. _____/20

4. _____/20

Total _____/100

Answer the following 4 miscellaneous questions about entropy and collision theory.

1a. (5 pts). What is the statistical prediction for the entropy of a perfect crystal at $T=0$? Give a mathematical explanation.

There is only one way to order a crystal, so $W=1$ and $S = k \ln W = 0$.

1b. (5 pts.) If there are two ways to orient each molecule in the crystal from 0a, how many ways W are there to orient a mole (N_A - Avagardo's number) of such molecules? Use that result to calculate the entropy of the crystal of N_A molecules.

$$\begin{aligned} W &= 2^{N_A} \\ S &= k \ln 2^{N_A} \\ &= k N_A \ln 2 \\ &= R \ln 2 \end{aligned}$$

1c (5 pts). Qualitatively explain how the entropy would be different for a gas of indistinguishable and non-interacting polyatomic versus monatomic molecules.

the rotational entropy would be higher due to the ^{2nd} _{3rd} moment of inertia and the additional vibrational modes.

1d. (5 pts.) Consider an elementary bimolecular reaction. If the masses of the gaseous molecules magically increased their masses by a factor of three (such as by forming non-reactive trimers), what does collision theory predict for the change in the rate constant?

$$k_2 = \sigma_c \langle v_{rel} \rangle N_A$$

$$\uparrow$$

$$\left(\frac{8kT}{\pi\mu}\right)^{1/2}$$

$$\mu = \frac{m_1 m_2}{m_1 + m_2}$$

$$= \frac{3m_1 \cdot 3m_2}{3m_1 + 3m_2}$$

$$= \frac{9}{3} \mu = 3\mu$$

Rate decreases by $\frac{1}{3}$

2. To estimate the population of a state, scientists often talk about energy levels in terms of kT . Answer the following questions to learn why.

2a. (5 pts.) Derive the equation for q^v for a harmonic oscillator.

$$q^v = \sum_i e^{-\beta \epsilon_i}$$

$$= 1 + \underbrace{e^{-\beta \epsilon}}_x + e^{-2\beta \epsilon} + e^{-3\beta \epsilon} \dots$$

$$= 1 + x + x^2 + x^3 \dots$$

$$= \frac{1}{1-x} = \frac{1}{1-e^{-\beta \epsilon}}$$

not necessary for exam
↓

$$S = 1 + x + x^2 \dots$$

$$xS = x + x^2 + x^3 \dots = S - 1$$

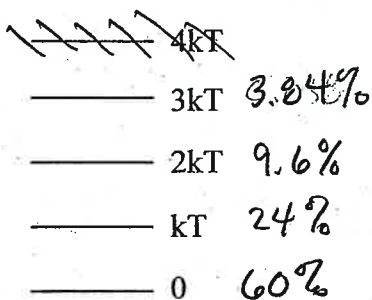
$$\therefore 1 = S - xS = S(1-x)$$

$$S = \frac{1}{1-x}$$

2b. (5 pts.) Write an expression that gives the population of the eigenstate ϵ_i for a harmonic oscillator.

$$P_i = \frac{n_i}{g^v} = \frac{e^{-\beta \epsilon_i}}{\frac{1}{1-e^{-\beta \epsilon}}} = e^{-\beta \epsilon_i} (1-e^{-\beta \epsilon})$$

2c. (10 pts) Consider a harmonic oscillator that has an energy level spacing of kT . For each eigenstate drawn below, label its population. In %.



$$q = (1 - e^{-\beta \epsilon})^{-1} = (1 - e^{-\epsilon/kT})^{-1} = (1 - e^{-1})^{-1} = 1 - 0.4 = 0.6$$

$$P_0 = e^0 \cdot 0.6 = 0.6$$

$$P_1 = e^{-1} \cdot 0.6 = 0.4 \cdot 0.6 = 0.24$$

$$P_2 = e^{-2} \cdot 0.6 = (e^{-1})^2 \cdot 0.6 = 0.4 \cdot 0.4 \cdot 0.6 = 0.096$$

$$P_3 = e^{-3} \cdot 0.6 = (0.4)^3 \cdot 0.6 = 0.096 \cdot 0.4 = 0.0384$$

2d. (5 pts) In lecture, we found that I_2 , which has a frequency of $\tilde{\nu} = 215 \text{ cm}^{-1}$, has 23% of its population in the $v=1$ level and that Cl_2 , which has a frequency of $\tilde{\nu} = 560 \text{ cm}^{-1}$, has 6% of its population in the $v=1$ level. Using your results above, what is the approximate population of $v=1$ for BrI , which has a frequency of $\tilde{\nu} = 410 \text{ cm}^{-1}$? Show your work.

exact BrI $\tilde{\nu} = 420 \approx 2kT$

$$q = (1 - e^{-2})^{-1} = (1 - 0.4 \cdot 0.4)^{-1} \approx 0.8$$

$$P_1 = e^{-2} \cdot 0.8 \approx 0.16 \times 0.8 \approx 0.2 \times 0.8 \approx 0.16 = 16\%$$

Notice from I_2 & Cl_2 that population can be estimated by comparing to kT .

if $I_2 = 215 \text{ cm}^{-1} \approx kT$ from above $P_1 = 24\%$ versus 23% in class.

Cl_2 is not quite $3kT$: from above $P_3 = 3.8\%$.

\therefore for $BrI = 410 \text{ cm}^{-1} \approx 2kT$; estimate $P = 16\%$.

Answer the following 2-part question to learn about the density of the atmosphere.

3a. (5 pts.) Write the expression for the ratio of the number of molecules in a state with energy ϵ_1 to the number in a state with energy ϵ_0 at a temperature T .

$$P_{\epsilon_1} = \frac{n_1}{g} = \frac{e^{-\beta\epsilon_1}}{g}$$

$$P_{\epsilon_0} = \frac{n_0}{g} = \frac{e^{-\beta\epsilon_0}}{g}$$

$$\frac{P_{\epsilon_1}}{P_{\epsilon_0}} = \frac{e^{-\beta\epsilon_1}}{e^{-\beta\epsilon_0}} = e^{-\beta(\epsilon_1 - \epsilon_0)}$$

3b. (10 pts.) The energy of a particle in a gravitational field is $\epsilon = mgh$ where m is the mass of the particle, h is the height above a reference point, and $g \approx 10 \text{ m/s}^2$ is the acceleration of gravity. Estimate the ratio of the number density of O_2 at 8314 m compared to that at sea level. Assume that the temperature is 320 K.

$$e^{-\beta(\epsilon_1 - \epsilon_0)} = \exp \left[- \frac{32 \text{ g/mole} \cdot \frac{\text{kg}}{1000 \text{ g}} \times \frac{10 \text{ m}}{\text{s}^2} (8314 \text{ m} - 0 \text{ m})}{\frac{8.314 \text{ J}}{\text{mol K}} \cdot 320 \text{ K}} \right]$$

$$= \exp \left[-10^{-1} \times 10^{-3} \times 10 \times 10^3 \right] = \exp[-1] = 0.4$$

$$\frac{P_{\epsilon_1}}{P_{\epsilon_0}} = e^{-1} = 0.4$$

so, pressure of O_2 at 8314 m is about 40% of that at sea level.

4a. (5 pts.) Derive the equation for the equilibrium constant K for the chemical reaction $A \leftrightarrow B$, where the eigenstates of A and B are spaced by ϵ_A and ϵ_B , respectively, and the relative energy of B is offset by $\Delta\epsilon_{AB}$. Start from the definition of $K = [B]/[A]$.

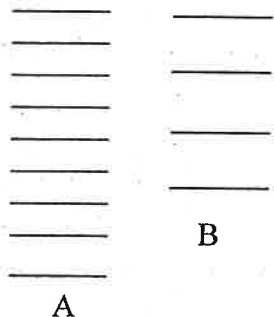
$$K = \frac{[B]}{[A]} = \frac{N_B}{N_A} \leftarrow \text{fraction in B}$$

fraction in A: $\frac{\# \text{ in A}}{\text{total}} = \frac{\sum_j e^{-\beta \epsilon_{jA}}}{\mathcal{Z}} = \frac{\mathcal{Z}_A}{\mathcal{Z}}$

$$\frac{\# \text{ in B}}{\text{total}} = \frac{e^{-\beta \Delta \epsilon_{AB}} \sum_k e^{-\beta \epsilon_{kB}}}{\mathcal{Z}} = \frac{\mathcal{Z}_B}{\mathcal{Z}}$$

$$K = \frac{N_B}{N_A} = \frac{\mathcal{Z}_B}{\mathcal{Z}_A} e^{-\beta \Delta \epsilon_{AB}} = \frac{\sum_k e^{-\beta \epsilon_{kB}}}{\sum_j e^{-\beta \epsilon_{jA}}} e^{-\beta \Delta \epsilon_{AB}}$$

4b. (5 pts) If the eigenstates have the relative spacings drawn below, explain whether the equilibrium constant will favor A or B and why.



Will favor A because of the more closely spaced eigenstates and because of the lower grand state.

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4c. (5 pts) Calculate K for $\epsilon_A = kT$, $\epsilon_B = 2kT$ and $\Delta\epsilon_{AB} = 2kT$ for vibrational mode in the high temperature limit. Round your answer to 1 significant digit.

high temp limit, doesn't ask question

$$K = \frac{g_B}{g_A} e^{-\beta\Delta\epsilon_{AB}} = \frac{(1 - e^{-\beta\epsilon_A})}{(1 - e^{-\beta\epsilon_B})} e^{-\beta\Delta\epsilon_{AB}}$$

$$= \frac{(1 - e^{-1})}{(1 - e^{-2})} e^{-2}$$

$$= \frac{(1 - 0.4)}{(1 - 0.16)} \cdot 0.16$$

ok without

high temp limit, could
Taylor expand $1 - e^{-\beta\epsilon} = 1 - 1 + \beta\epsilon - \dots$

$$\frac{\epsilon_A}{\epsilon_B} \cdot e^{-\beta\Delta\epsilon_{AB}} = \frac{1}{2} \cdot 0.16$$

$$= 0.08 \approx 0.1$$

$$= \frac{0.6}{0.84} \cdot 0.16 \approx \frac{0.6}{0.8} \times 0.2 = \frac{0.6}{4} \approx 0.15$$

4d. (5 pts) To make $K = 1$, what value would $\Delta\epsilon_{AB}$ need to be? Write an expression (do not evaluate).

$$1 = \frac{g_B}{g_A} e^{-\beta\Delta\epsilon_0}$$

$$e^{-\beta\Delta\epsilon_0} = \frac{g_A}{g_B}$$

$$-\beta\Delta\epsilon_0 = \ln \frac{g_A}{g_B}$$

$$\Delta\epsilon_0 = + \frac{1}{\beta} \ln \frac{g_B}{g_A}$$

5a. (5 pts) Write out the equation for the canonical partition function Q in terms of the molecular translational, rotational, and vibrational partition functions (e.g. q^T , q^R , q^V) for a homonuclear diatomic gas in the high temperature limit.

$$Q = q_T q_R q_V \quad Q = \frac{q^N}{N!} = \frac{(q_T q_R q_V)^N}{N!}$$

from the ~~ba~~ eqn page

$$Q = \frac{1}{N!} \left[\frac{V}{\Lambda^3} \times \frac{1}{\sigma} \frac{kT}{hcB} \times \frac{1}{(1 - e^{-hc\tilde{\nu}/kT})} \right]^N$$

5b. (5 pts.) Find the equation for $\langle E^T \rangle$ for the homonuclear diatomic gas.

$$q^T = q^T_x q^T_y q^T_z = \frac{V}{\Lambda^3} = \frac{V}{\beta^{3/2} (\dots)}$$

$$\Lambda = \left(\frac{h^2 \beta}{2\pi m} \right)^{1/2} = \beta^{1/2} (\dots)$$

$$\langle E^T \rangle = - \left(\frac{\partial \ln q^T}{\partial \beta} \right)_V = + \frac{3}{2} \left(\frac{\partial \ln \beta}{\partial \beta} \right)_V$$

$$= \frac{3}{2} \frac{1}{\beta}$$

$$= \frac{3}{2} kT$$