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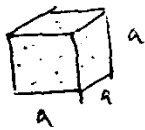
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Course 565/665 Lecture Number _____ Date 4/1/03

Lecturer Dr. Silvia Cavagnaro Note Taker Eric Fulmer

Last Time

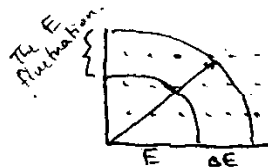


$$M = \frac{1}{8} \left(\frac{4}{3} \pi R^3 \right) = \frac{\pi}{6} \left(\frac{8m\epsilon}{h^2} \right)^{3/2} V$$

$$V = a^3$$

where M is the number of positive integers.

Towards Fluctuations of Energy:



$$W(E) \Delta E$$

↑
of positive integer points in the shell between the energies E and $(E + \Delta E)$.

$$W(E) \Delta E = M(E + \Delta E) - M(E)$$

$$\approx \frac{dM}{dE} \Delta E$$

$$W(E) \Delta E = \frac{\pi}{4} \left(\frac{8m}{h^2} \right)^{3/2} \cdot V E^{1/2} \cdot \Delta E$$

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Example An ensemble of particles in gaseous state

$$m = 40 \text{ g/mol} = 6.64 \times 10^{-26} \text{ kg/atom}$$

$$T = 300 \text{ K}$$

$$\begin{aligned} \epsilon &= \frac{3}{2} kT = \frac{3}{2} (1.38 \times 10^{-23} \text{ J/K}) (300 \text{ K}) \\ &= 6.21 \times 10^{-21} \text{ J} \end{aligned}$$

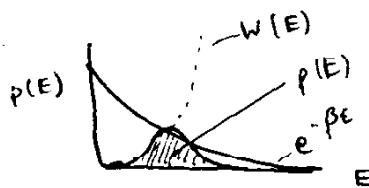
$$\Delta \epsilon = 0.01 \epsilon \quad \text{--- i.e. 1\% } \epsilon \text{ fluctuations.}$$

$$\begin{aligned} W(E) \Delta E &= \frac{\pi}{4} \left(\frac{8 \times 6.64 \times 10^{-26} \text{ kg/atom}}{(6.626 \times 10^{-34} \text{ J s})^2} \right) \times (10^{-2} \text{ m})^2 \\ &\times (6.21 \times 10^{-21} \text{ J})^{3/2} (0.01) = 5.11 \times 10^{24} \text{ states} \end{aligned}$$

Conclude: Even a small E fluctuation (1%) gives a large variation in the density of states. ($W(E) \Delta E$).

Energy fluctuations

$$p(E) = \frac{W(E) e^{-\beta E}}{Q}$$



$p(E)$ is maximized where both $W(E)$ and $e^{-\beta E}$ are similar in magnitude

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The shape of $p(E)$ is well approximated to be Gaussian.

$$\langle E \rangle = U$$

$$\ln p(E) = \ln p(U) + \left(\frac{\partial \ln p(E)}{\partial E} \right)_{E=U} (E-U) + \frac{1}{2} \left(\frac{\partial^2 \ln p(E)}{\partial E^2} \right)_{E=U} (E-U)^2 + \dots$$

$$T(E) \equiv T, \quad T = \left(\frac{\partial U}{\partial S} \right)_{V,N}$$

$$\ln [p(E)] = \ln \left(\frac{W(E)}{Q} \right) - \beta E$$

$$p(E) = \frac{W(E) e^{-\beta E}}{Q}$$

Q is a sum over E levels

$$S(E) = k \ln W(E); \quad \left(\frac{\partial S(E)}{\partial E} \right)_{V,N} = \frac{1}{T(E)}$$

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$$\frac{\partial \ln p(E)}{\partial E} = \frac{\partial \ln W(E)}{\partial E} - \beta = \frac{\partial (S/k)}{\partial E} - \beta$$

$$= \frac{1}{k} - \frac{1}{T(E)} - \frac{1}{kT_0}$$

$$C_v = \left(\frac{\partial E}{\partial T} \right)_v ; \quad \frac{\partial^2 \ln p(E)}{\partial E^2} = \frac{1}{kT^2} \left(\frac{\partial T}{\partial E} \right)_{E=U} = \frac{1}{kT_0^2 C_v}$$

$$p(E) = p(U) e^{-\frac{(E-U)^2}{2kT_0^2 C_v}}$$

$$p(E) = e^{[k - T_0 S(U)]} e^{-\frac{(E-U)^2}{2kT_0^2 C_v}}$$

This is a Gaussian Function.

Variance $\sigma^2 = \langle (x-a)^2 \rangle$

$$\sigma^2 = \langle (E-U)^2 \rangle = \langle E^2 \rangle - \langle E \rangle^2$$

$$= \langle E^2 \rangle - U^2 = kT_0^2 C_v$$

Standard Deviation: $\sigma = \sqrt{\sigma^2} = \sqrt{kT_0^2 C_v} \quad \sigma \sim \sqrt{C_v}$

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The width of the energy fluctuations is proportional to C_v .

C_v is extensive; $C_v \sim N$.

$$\frac{\sigma_E}{E} \sim \frac{\sqrt{N}}{N} = \frac{1}{\sqrt{N}}$$

As $N \rightarrow 10^{23}$, $\frac{1}{\sqrt{N}} \rightarrow 0$.
Thus, fluctuations become very small for molar amounts.

Do Example 12.2